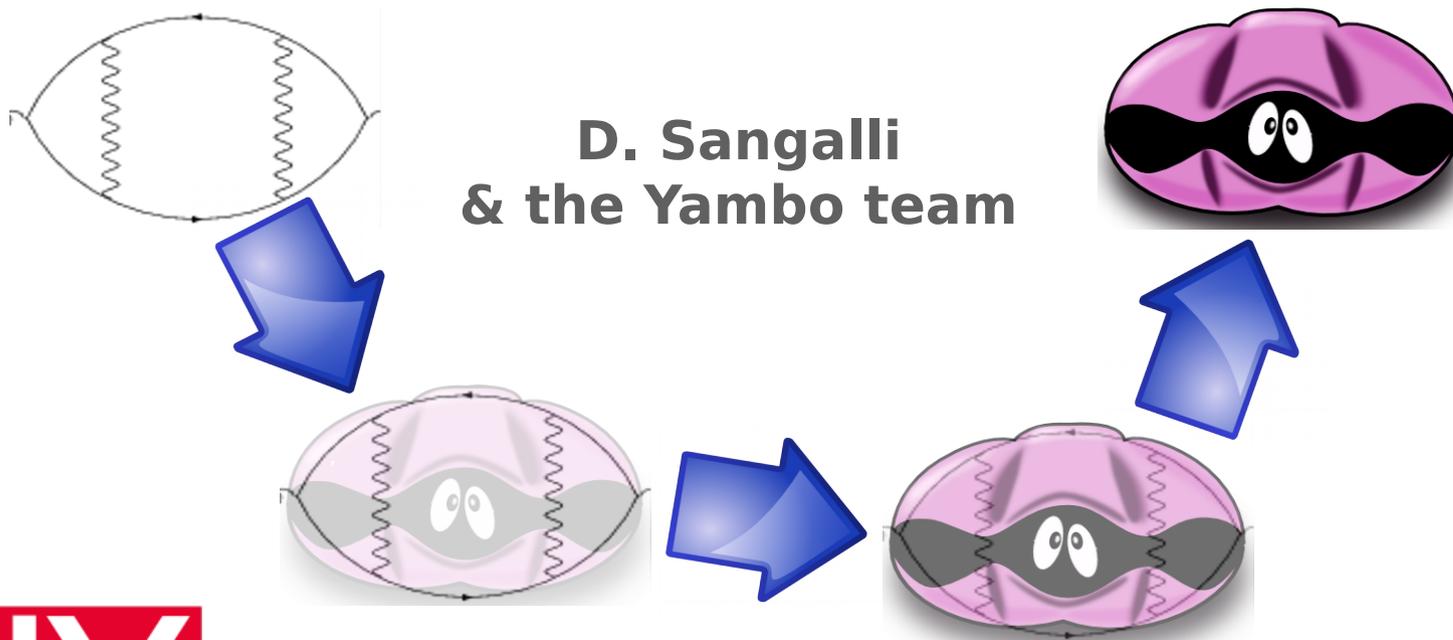




CNR
Istituto di Struttura
della Materia

CNR-ISM, Division of Ultrafast
Processes in Materials (FLASHit),
Area della Ricerca di Roma 1,
Monterotondo Scalo, Italy

Introduction to the Yambo code: Past, present and future



Bologna
4th Dec 2017



- 11.00-11.45 Introduction to the Yambo code: Past, Present and Future (D. Sangalli)
- [2] 11:45-12:30 Theoretical spectroscopy: GW and ARPES (A. Ferretti)
- [3] 14.00-14.45 Yambo parallelization: strategies & performances (Sangalli)
- [4] 14:45-15:30 HPC use cases: electronic and optical properties of graphene nano-ribbons (A. Ferretti)

outline

- ❑ A bit of history
- ❑ The code: **the structure of yambo**
- ❑ Applications: **few examples**
- ❑ Users view: **how it works in practice**
- ❑ Future: **new developments and more**

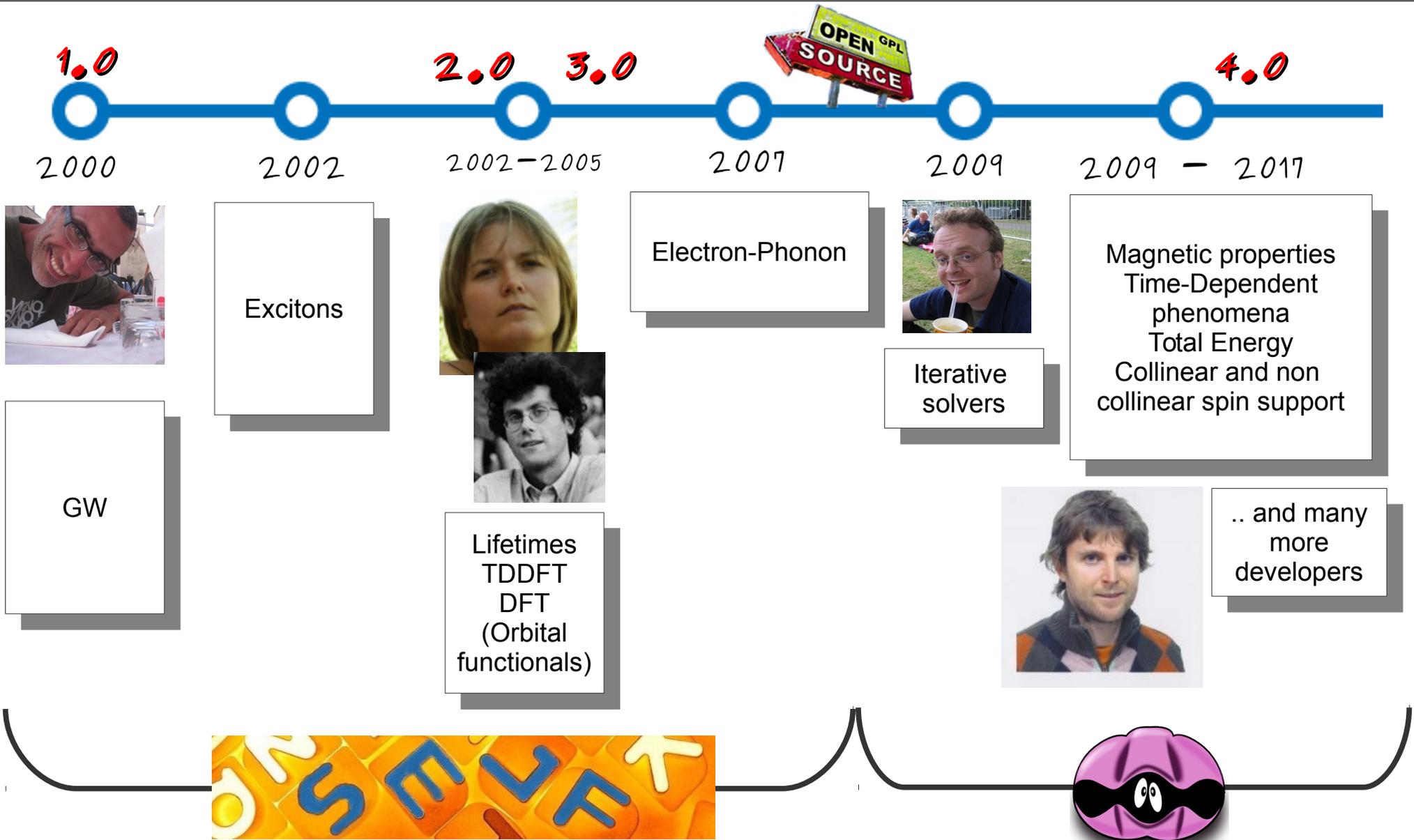
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A bit of history



A bit of history



Yambo: an ab initio tool for excited state calculations,
A. Marini, C. Hogan, M. Grüning, D. Varsano,
Comp. Phys. Comm. 180, 1392 (2009).

A new paper
coming soon ...

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What is yambo ?

Interfaces

Planewave

Pseudopotential codes:

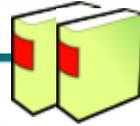
abinit.org



www.yambo-code.org

What is yambo ?

Theory



Many-Body perturbation Theory

Time-dependent density
functional theory

Interfaces

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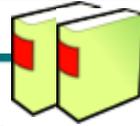
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Libraries

ScaLAPACK



libxc

FFTW

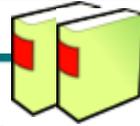


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Different projects



Developers



What can yambo do ?

Properties

GPL

Quasi-particles

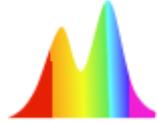
Optical absorption

Electron energy loss

Magneto optical properties

Electron-phonon coupling

Surface spectroscopy



DEVELOPMENT / PRE-GPL

Out of equilibrium dynamics

Multi-level parallelization



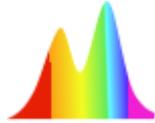
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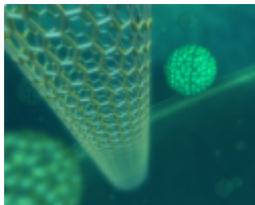
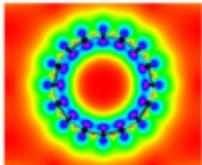
Quasi-particles
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DEVELOPMENT / PRE-GPL

Out of equilibrium dynamics
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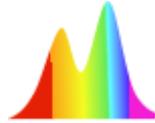
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What can yambo do ?

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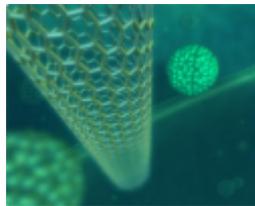
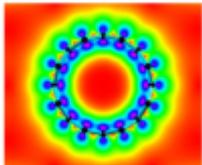
Quasi-particles
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DEVELOPMENT / PRE-GPL

Out of equilibrium dynamics
Multi-level parallelization

Applications



Community & Publications

Growing community of users using Yambo for fore-front research. More than 200 publications.



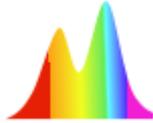
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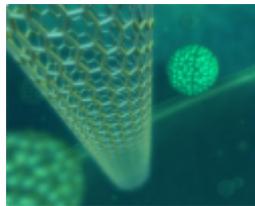
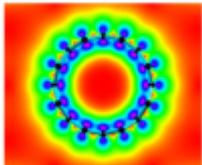
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DEVELOPMENT / PRE-GPL

Out of equilibrium dynamics
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Applications



Schools

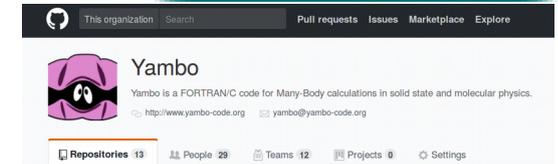


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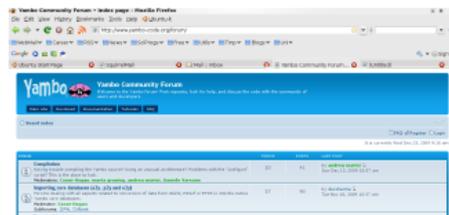


Support & reach out



Git repository

www.yambo-code.org



Dedicated User Forum

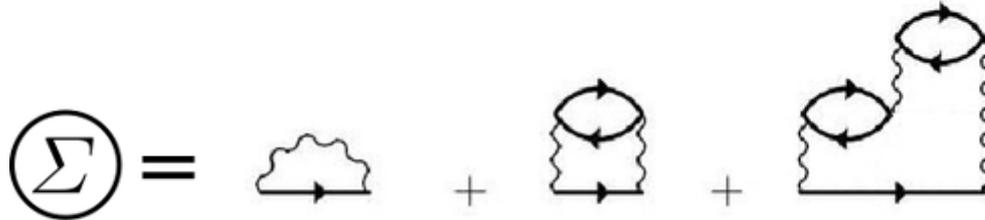
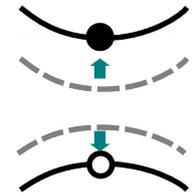


Online documentation and tutorials

Basic ingredients

Quasi-Particle energies (HF, Cohsex, GW)

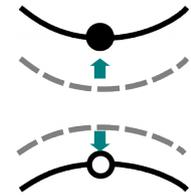
$$\epsilon_{nk}^{QP} = \epsilon_{nk}^{KS} + \langle \psi_{nk} | \Sigma(\epsilon_{nk}^{QP}) - V_{xc} | \psi_{nk} \rangle$$



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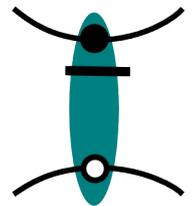
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Absorption spectra (TDDFT, TDHF, BSE)

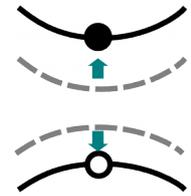
$$\epsilon(\omega) \propto \sum_I \sum_{nm, n'm'} \sum_{kk'} \frac{(x_{nmk} A_{nmk}^I)(A_{m'n'k'}^I x_{n'm'k'})^*}{\omega - \omega_I + i\eta}$$



Basic ingredients

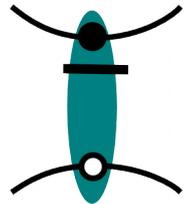
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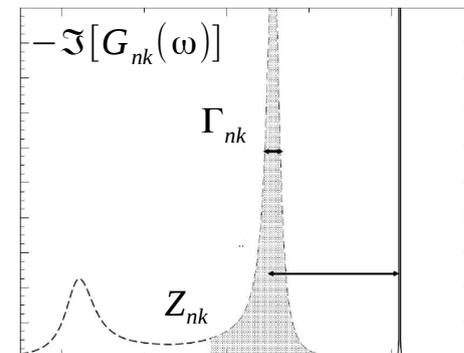
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More: QP lifetimes and QP weights, Satellites

$$\Gamma_{nk} = \Im[\Sigma_{nk}] \quad Z_{nk} = \left[1 - \frac{\partial \Sigma_{nk}(\omega)}{\partial \omega} \right]^{-1}$$

$$\Im[G_{nk}(\omega)] = \Im[G_{nk}^0(\omega) - \Sigma_{nk}(\omega)]^{-1}$$



outline

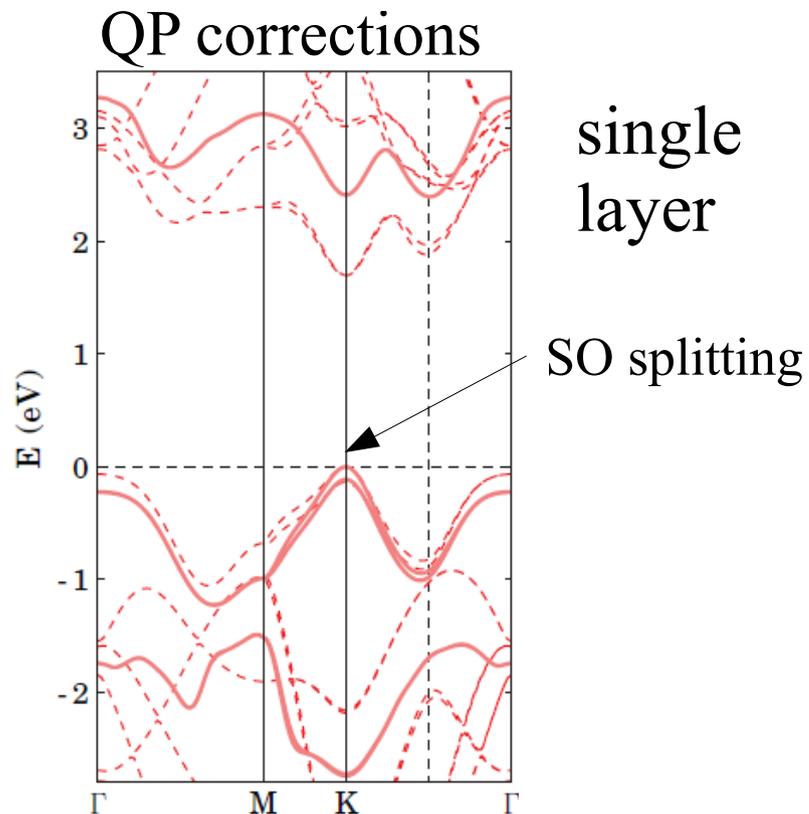
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QP and BSE with SOC

PHYSICAL REVIEW B 88, 045412 (2013)

Effect of spin-orbit interaction on the optical spectra of single-layer, double-layer, and bulk MoS₂

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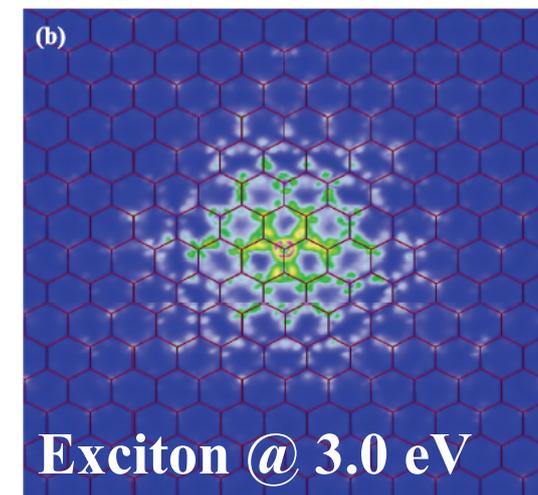
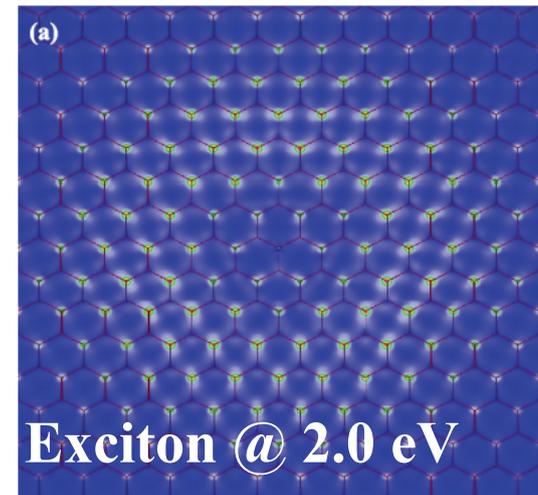
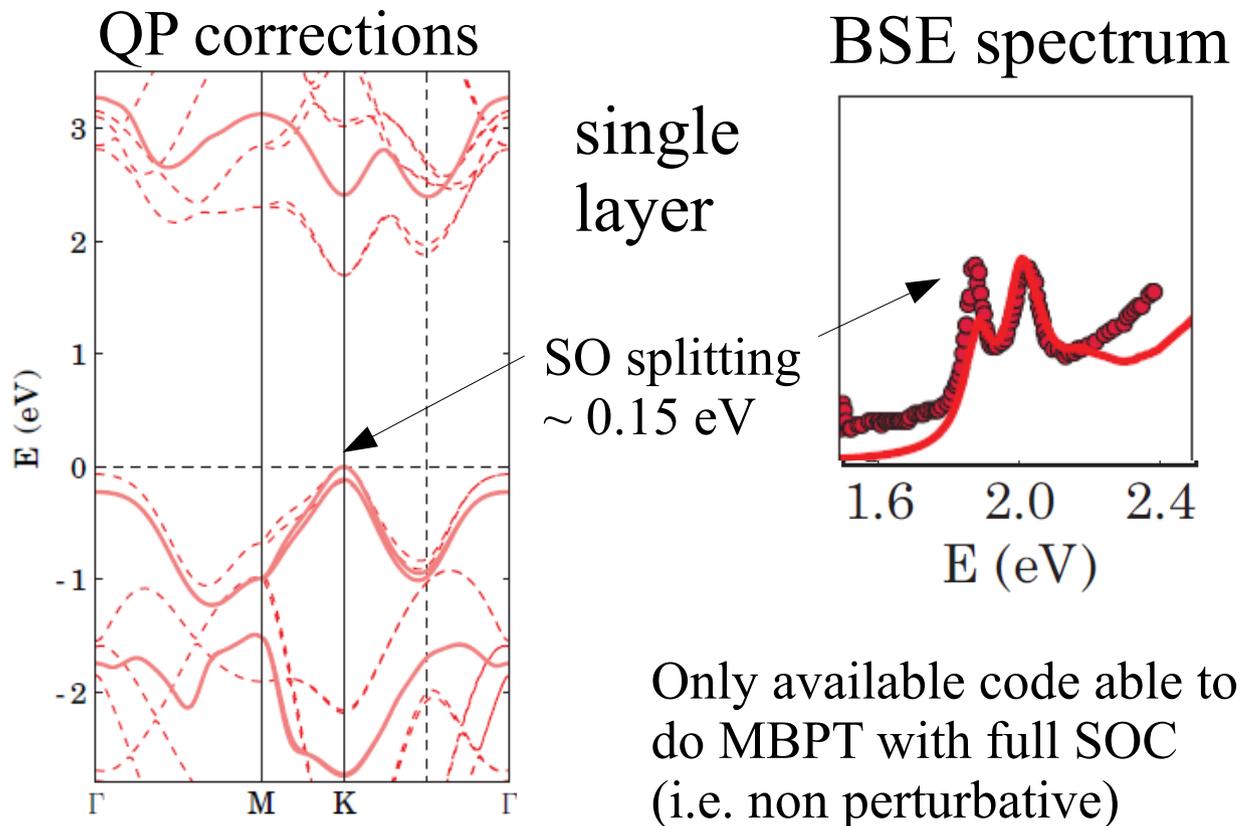


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Lifetimes in nanostructures

ARTICLES

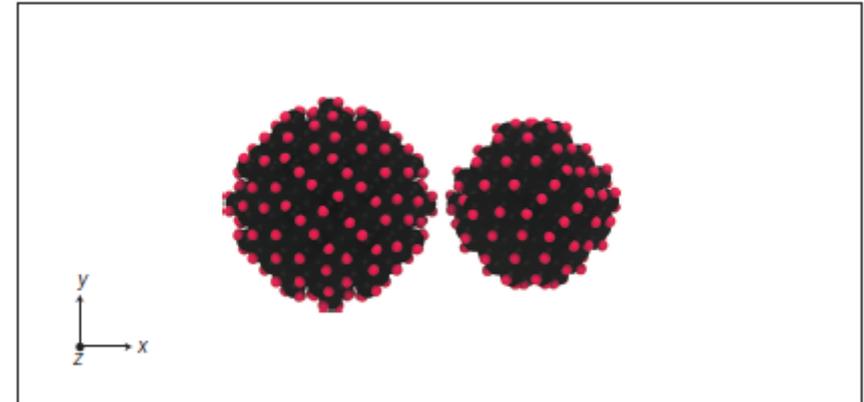
PUBLISHED ONLINE: 9 SEPTEMBER 2012 | DOI: 10.1038/NPHOTON.2012.206

nature
photonics

Silicon Nano-Crystals
440 Si atoms + 272 H atomsc

Carrier multiplication between interacting nanocrystals for fostering silicon-based photovoltaics

Marco Govoni^{1*}, Ivan Marri^{2*} and Stefano Ossicini^{2,3}



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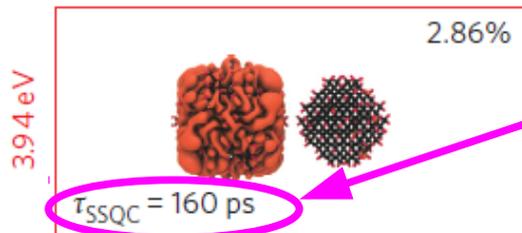
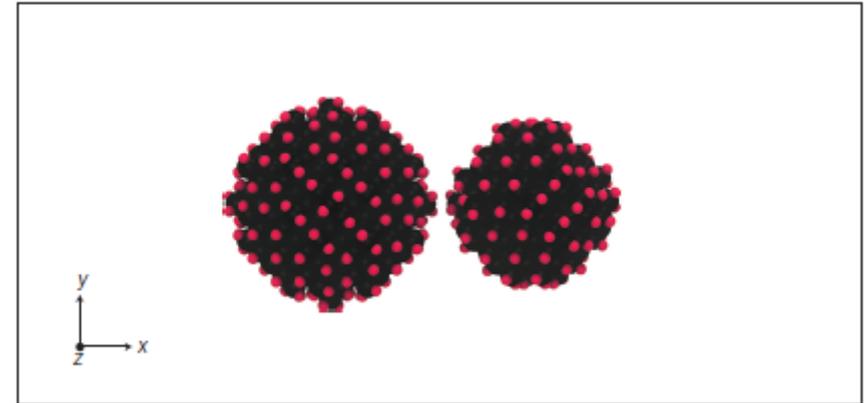
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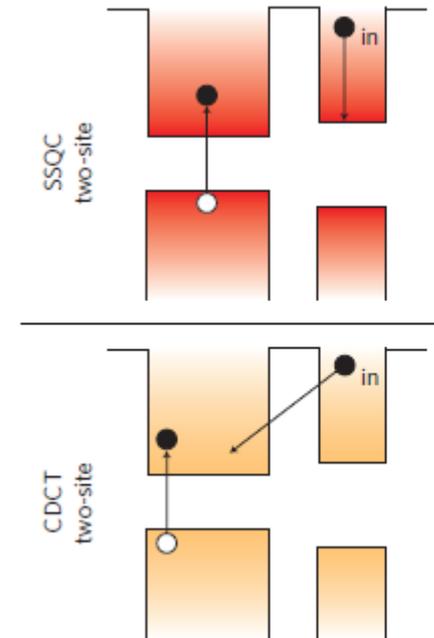
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Carriers multiplications
lifetimes



Lifetimes in nanostructures

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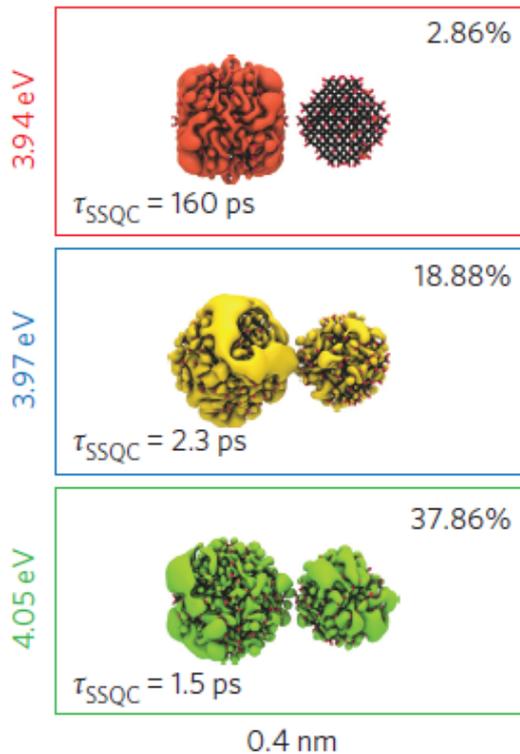
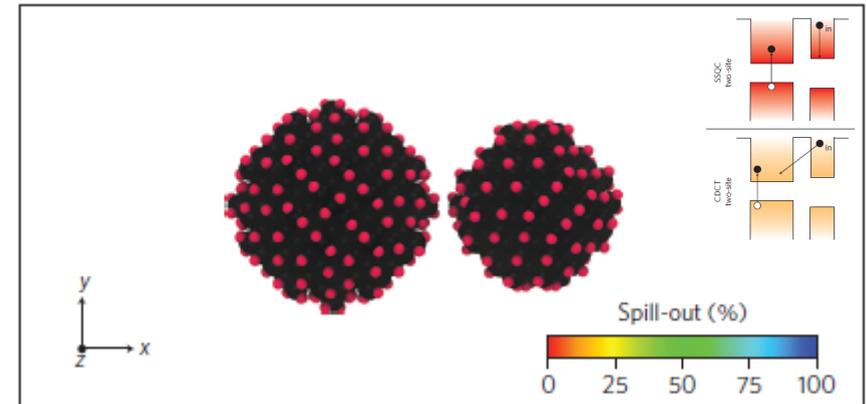
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Carriers multiplications
lifetimes

As a function of energy

Lifetimes in nanostructures

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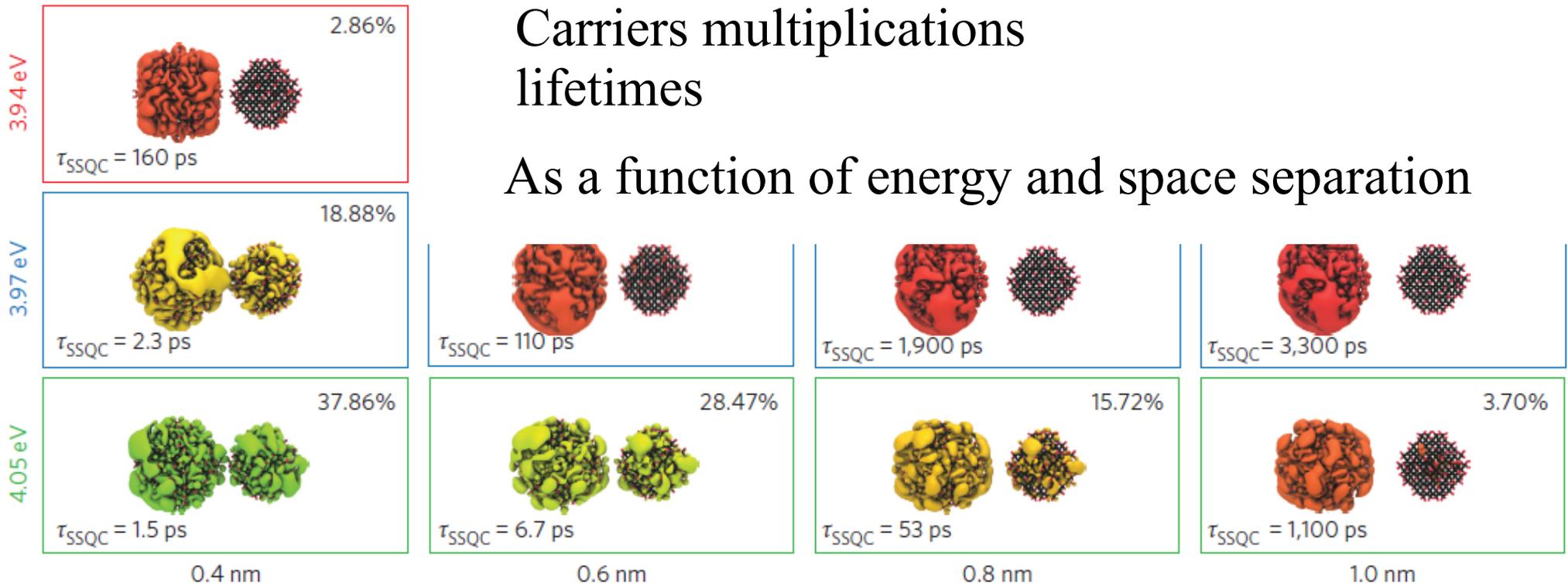
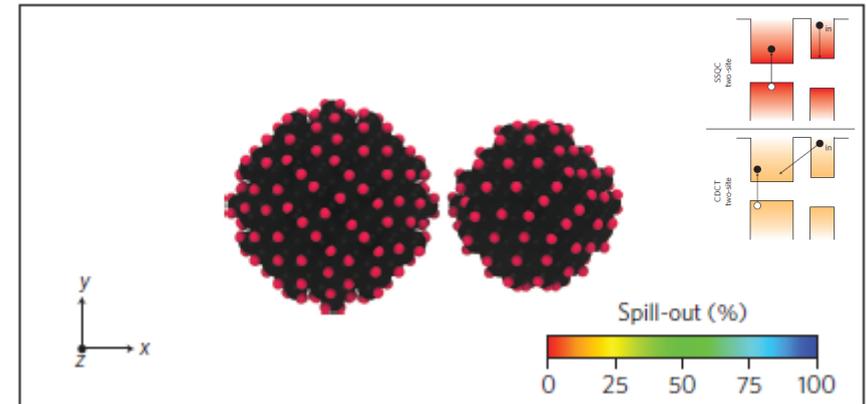
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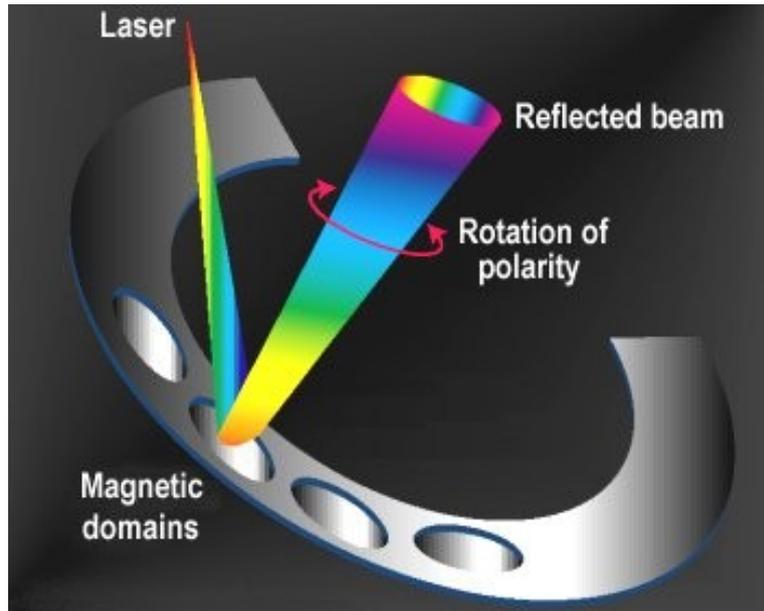
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More: magneto-optics

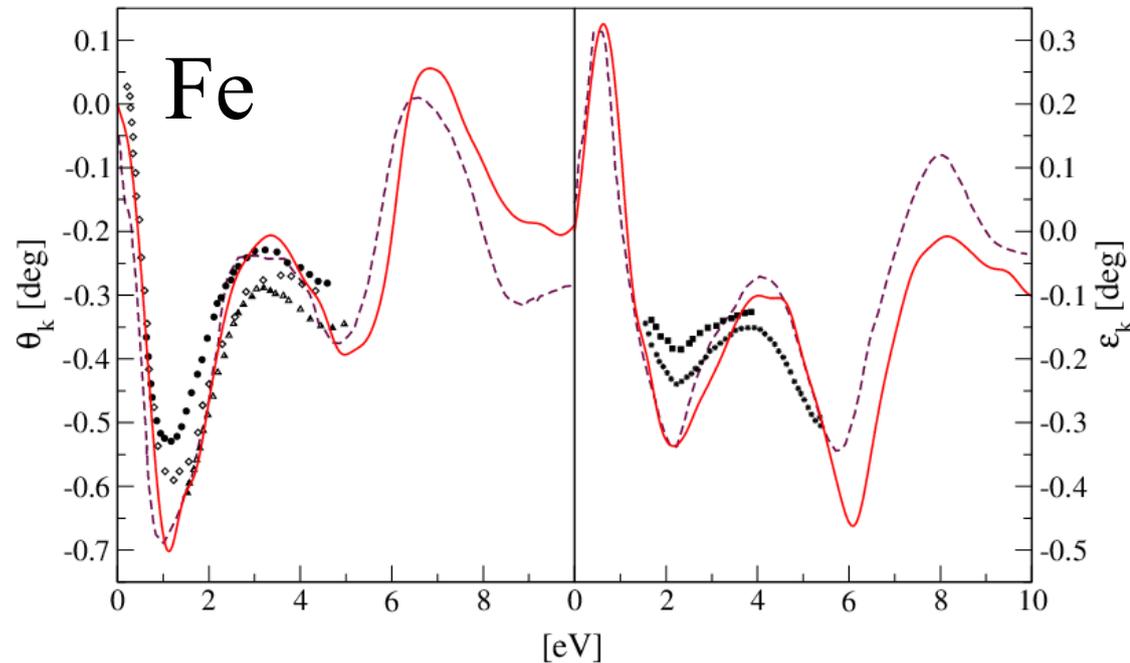
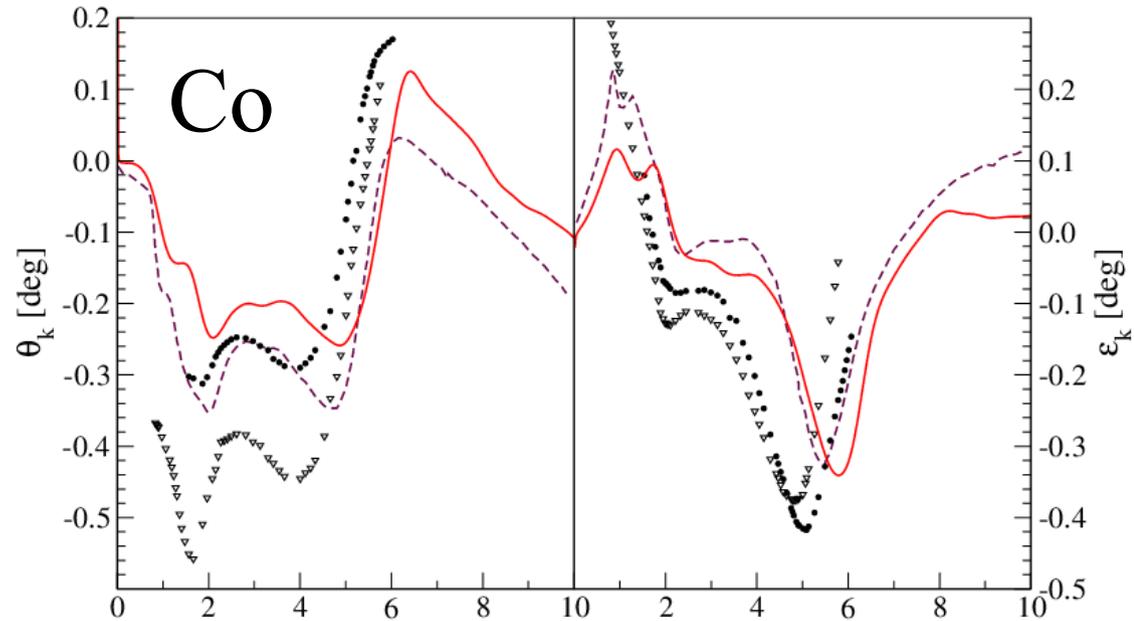


D. Sangalli et al. Phys. Rev. **B 86**, 125139 (2012)

Need the off-diagonal dielectric function

$$\epsilon_{\pm}(\omega) = \epsilon_{xx}(\omega) \pm i\epsilon_{xy}(\omega)$$

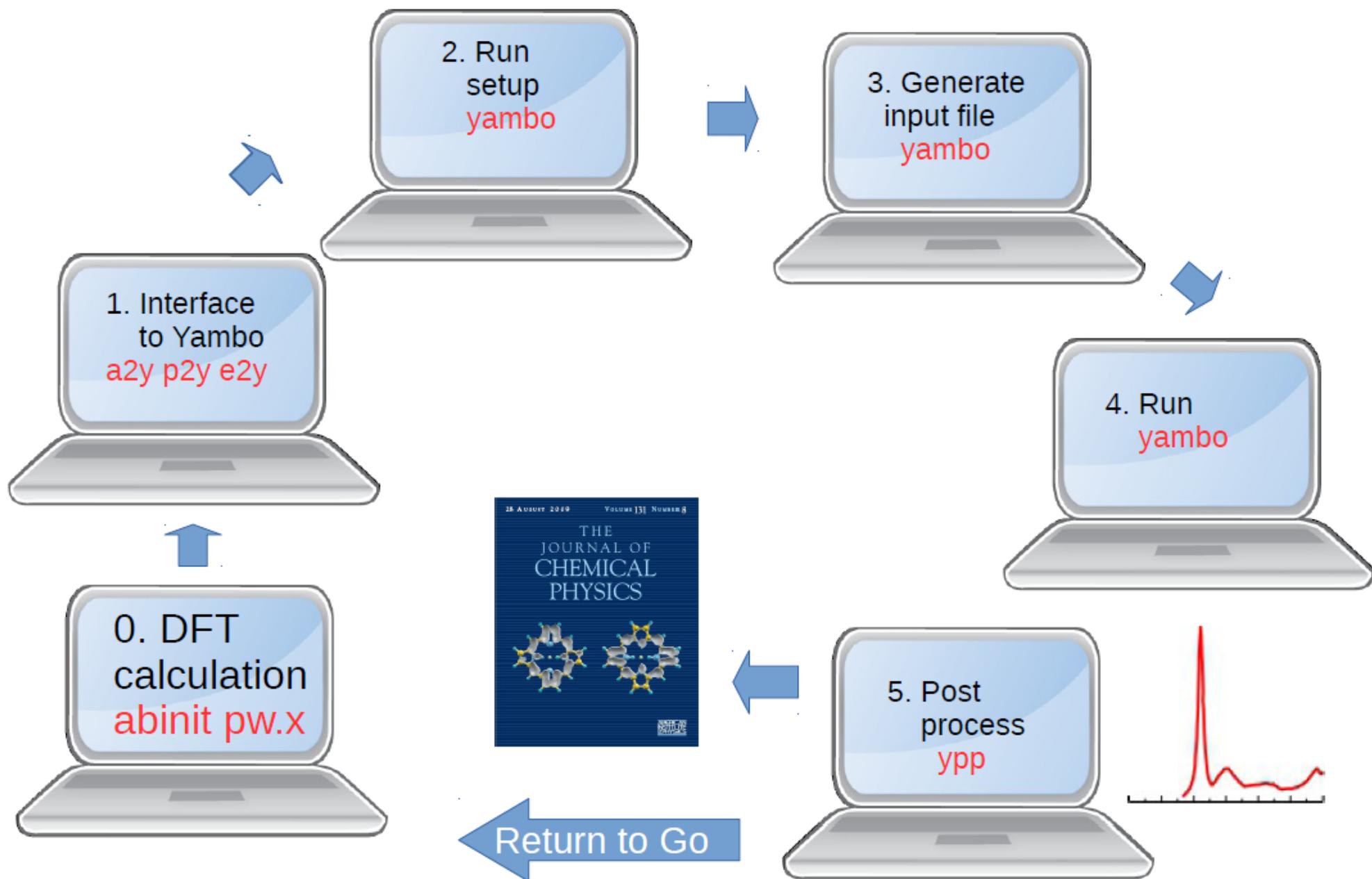
Needs to fully account for the spinorial wave-functions



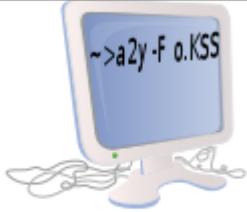
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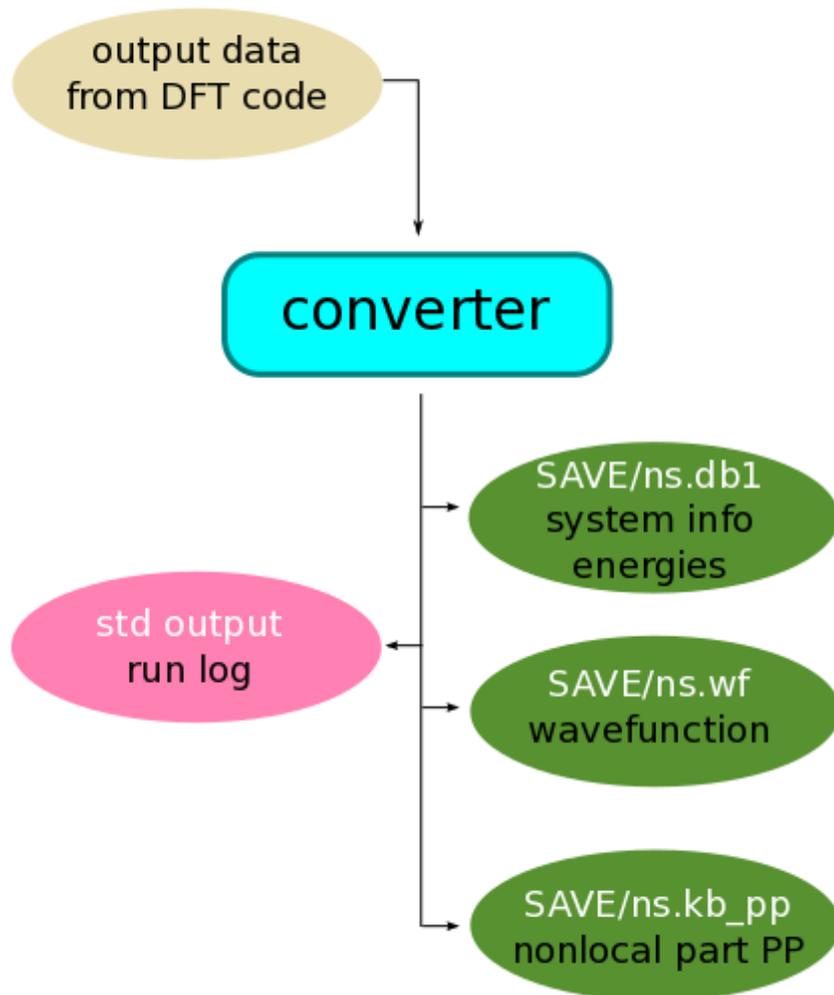
How it works



How it works



1. Generate the core databases



example: convert abinit KSS file with a2y

```
% ls  
3D_LiF_o_DS2_KSS
```

```
% a2y -F 3D_LiF_o_DS2_KSS
```

```
<--> [01] A(binit) 2 Y(ambo)  
<--> Checking input file ...3D_LiF_o_DS2_KSS  
<--> DBs path set to :.  
<--> KSS Header...abinit version 5.8.4  
<--> :: Atoms positions...done  
<--> Symmetries...[ID yes]...[INV no]...[TR yes]...[SI yes]  
<--> RL vectors...done  
<--> Energies...done  
<--> Report:  
<--> :: Electrons : 8.000000  
<--> :: Temperature [ev]: 0.000000  
<--> :: Lattice factors [a.u.]: 7.703475 7.703475 7.703475  
<--> :: K-points : 10  
<--> :: Bands : 10  
<--> :: Spinor components : 1  
<--> :: Spin polarizations : 1  
<--> :: Spin orbit coupling : no  
<--> :: Symmetries [spatial]: 24  
<--> :: [T-rev]: yes  
<--> :: Max WF components : 1885  
<--> :: RL vectors (WF): 1885  
<--> :: RL vectors (CHARGE): 1885  
<--> :: XC potential : Perdew & Wang (xc)  
<--> :: Atomic species : 2  
<--> :: Max atoms/species : 1  
<--> == DB1 ...done ==  
<--> == DB2 + nlPP ...done ==
```

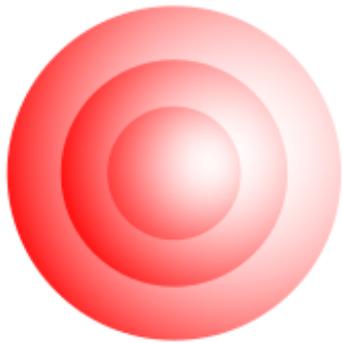
```
% ls SAVE  
ns.db1 ns.kb_pp ns.wf
```

How it works



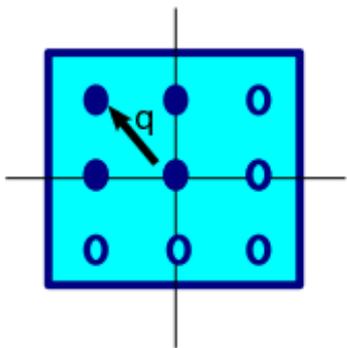
2. Run setup

= prepare general purpose databases for later use



* **Data initialization:**

reorders G-vectors into spherical shells
calculates Fermi level and electronic occupations
sets up energy grids



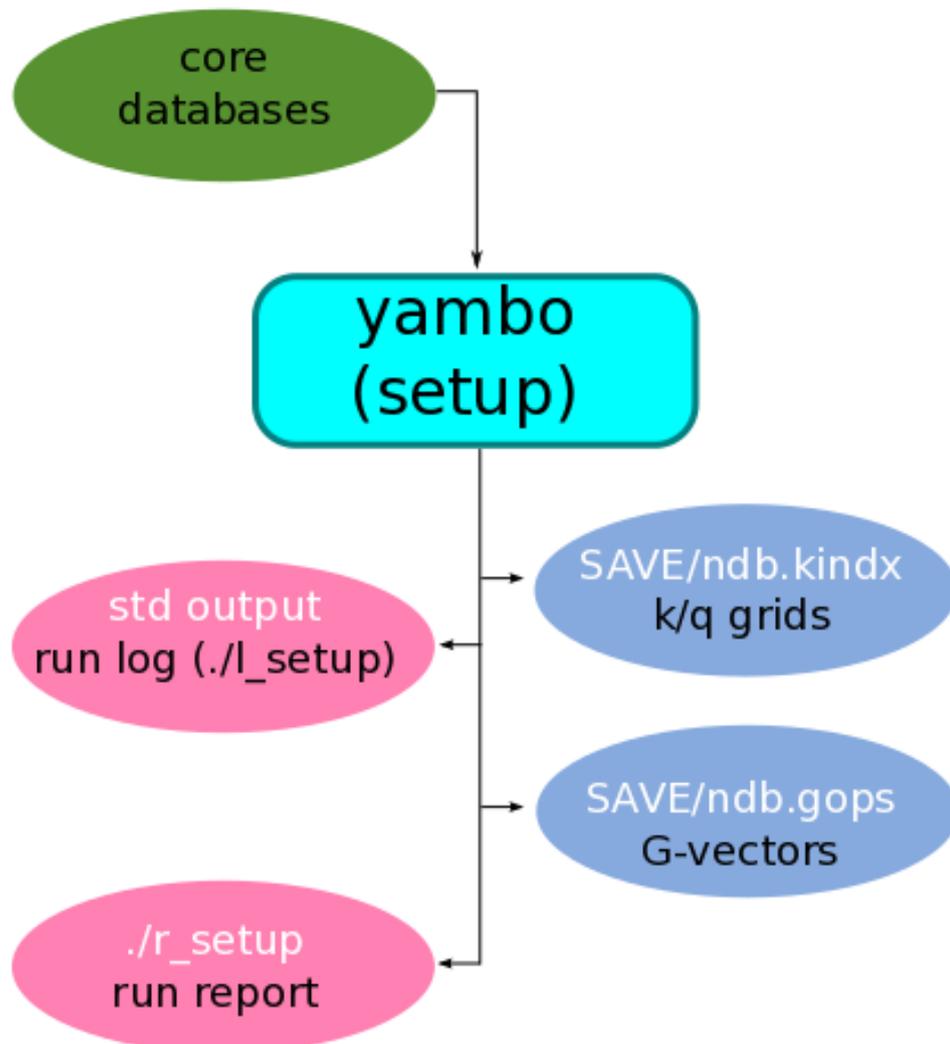
* **Brillouin-zone sampling:**

expands k-points to full BZ
generates q-point meshes
checks on uniformity of grids

How it works



2. Run setup



example: setup run

```
% ls SAVE/  
ns.db1 ns.kb_pp ns.wf
```

```
% yambo
```

```
<---> [01] Job Setup  
<---> [02] Input variables setup  
<---> [02.01] Unit cells  
<---> [02.02] Symmetries  
<---> [02.03] RL shells  
<---> Shells finder |#####| [100%] --(E) --(X)  
<---> [02.04] K-grid lattice  
<---> [02.05] Energies [ev] & Occupations  
<---> [03] Transferred momenta grid  
<---> X indexes |#####| [100%] --(E) --(X)  
<---> SE indexes |#####| [100%] --(E) --(X)  
<---> [04] Game Over & Game summary
```

```
% ls
```

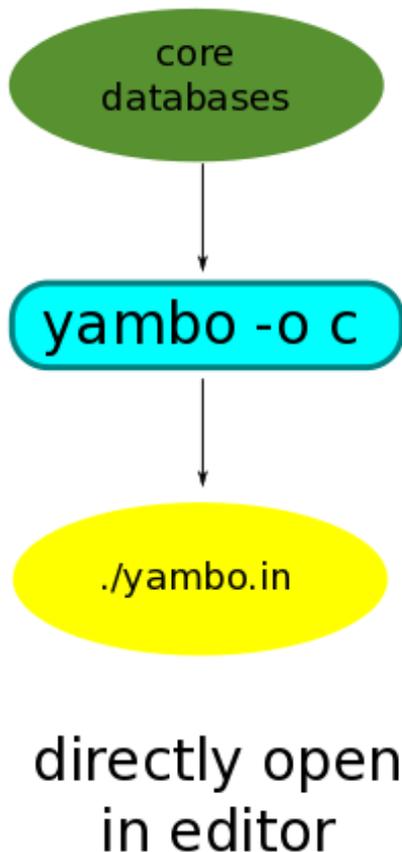
```
r_setup SAVE  
% ls SAVE  
ndb.gops ndb.kindx ns.db1 ns.kb_pp ns.wf
```

How it works



3. Generate input file

example: linear response (rpa level)



```
optics          # [R OPT] Optics
chi             # [R CHI] Dyson equation for Chi.
% QpntsRXd     # [Xd] Transferred momenta
  1 | 19 |
%
% BndsRnXd     # [Xd] Polarization function bands
  1 | 10 |
%
NGsBlkXd= 1    RL # [Xd] Response block size
% EnRngeXd     # [Xd] Energy range
  0.00000 | 10.00000 | eV
%
% DmRngeXd     # [Xd] Damping range
  0.10000 | 0.10000 | eV
%
ETStpsXd= 100  # [Xd] Total Energy steps
% LongDrXd     # [Xd] [cc] Electric Field
  1.000000 | 0.000000 | 0.000000 |
%
```

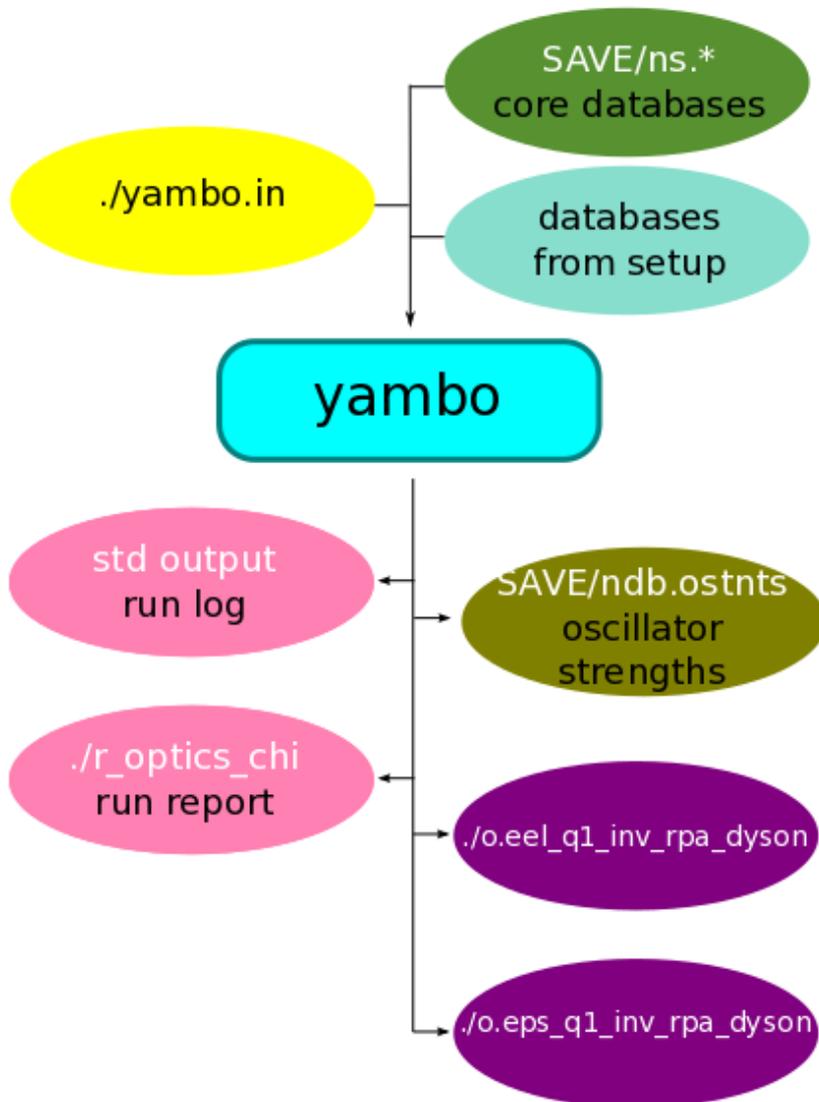
runlevels

default values
from existing
databases
(compatibility)

How it works



4. Run yambo



example: linear response (RPA)

```
% ls  
r_setup SAVE yambo.in
```

```
% yambo
```

```
<---> [01] Job Setup  
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<---> [02.01] Unit cells  
<---> [02.02] Symmetries  
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<---> [02.05] Energies [ev] & Occupations  
<---> [03] Transferred momenta grid  
<---> [04] Optics  
<---> [WF-Oscillators/G space loader] Wfs (re)loading | [...]  
<---> Dipole (T): |#####| [100%] --(E) --(X)  
<---> [FFT-X] Mesh size: 18 18 18  
<---> [WF-X loader] Wfs (re)loading |#####| [100%] --(E) --(X)  
<---> [X-CG] R(p) Tot o/o(of R) : 222 6144 100  
<---> [X] Upper matrix triangle filled  
<---> Xo@q[1] 1-100 |#####| [100%] --(E) --(X)  
<---> X @q[1] 1-100 |#####| [100%] --(E) --(X)  
<---> [05] Game Over & Game summary
```

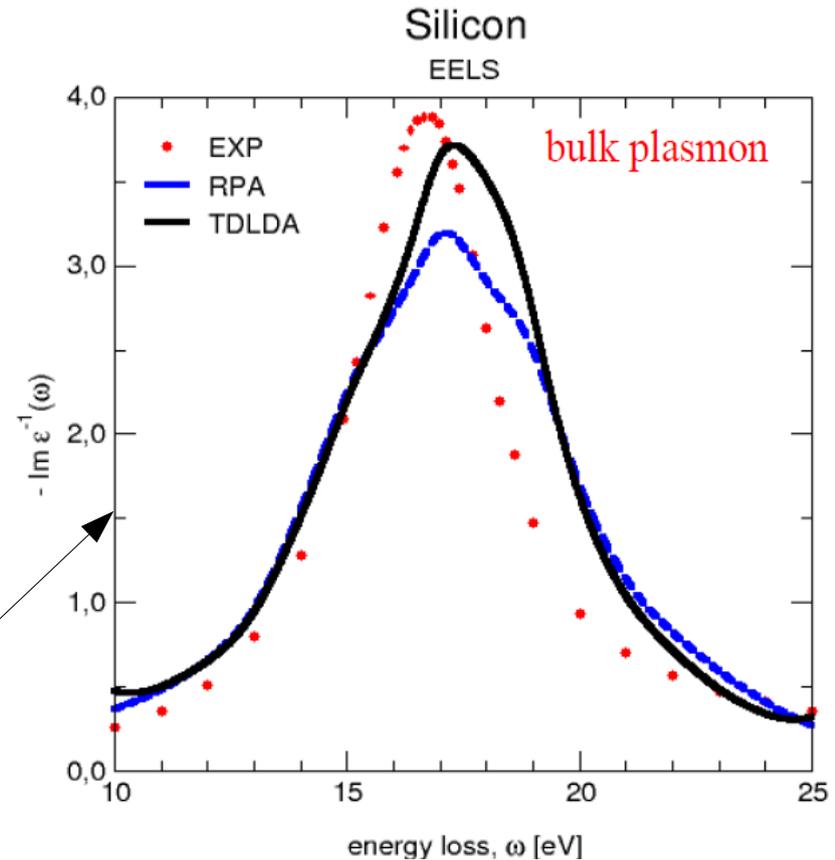
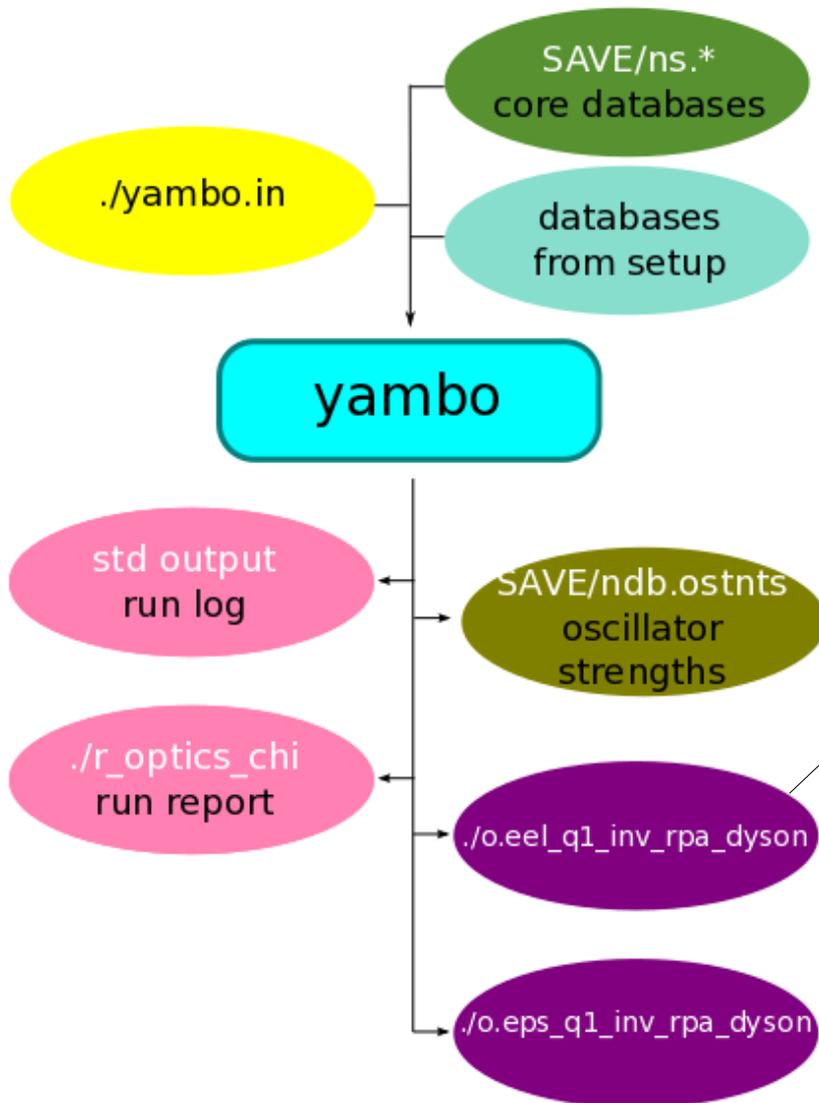
```
% ls  
o.eel_q1_inv_rpa_dyson o.eps_q1_inv_rpa_dyson  
r_optics_chi r_setup SAVE yambo.in
```

```
% ls SAVE  
ndb.gops ndb.kindx ndb.ostnts ns.db1 ns.kb_pp ns.wf
```

How it works



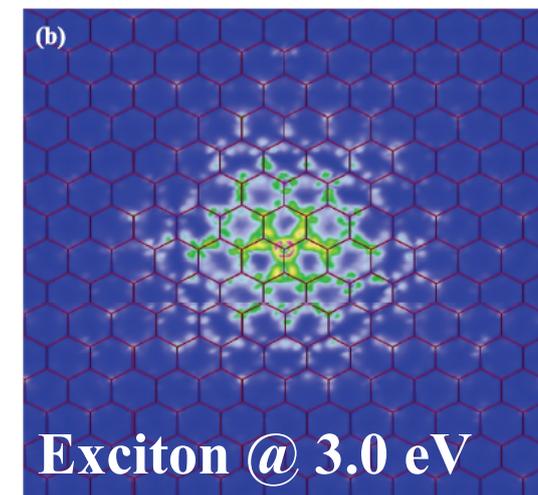
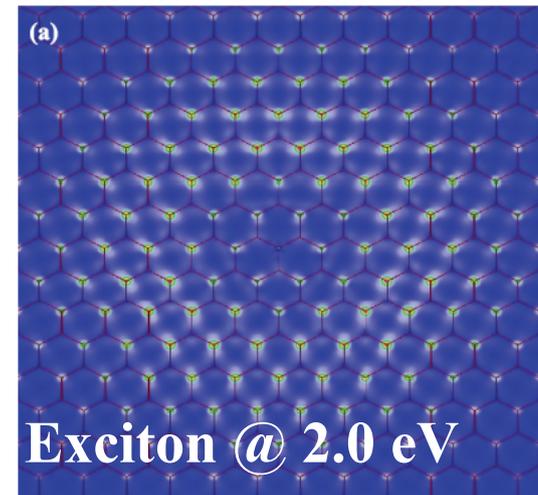
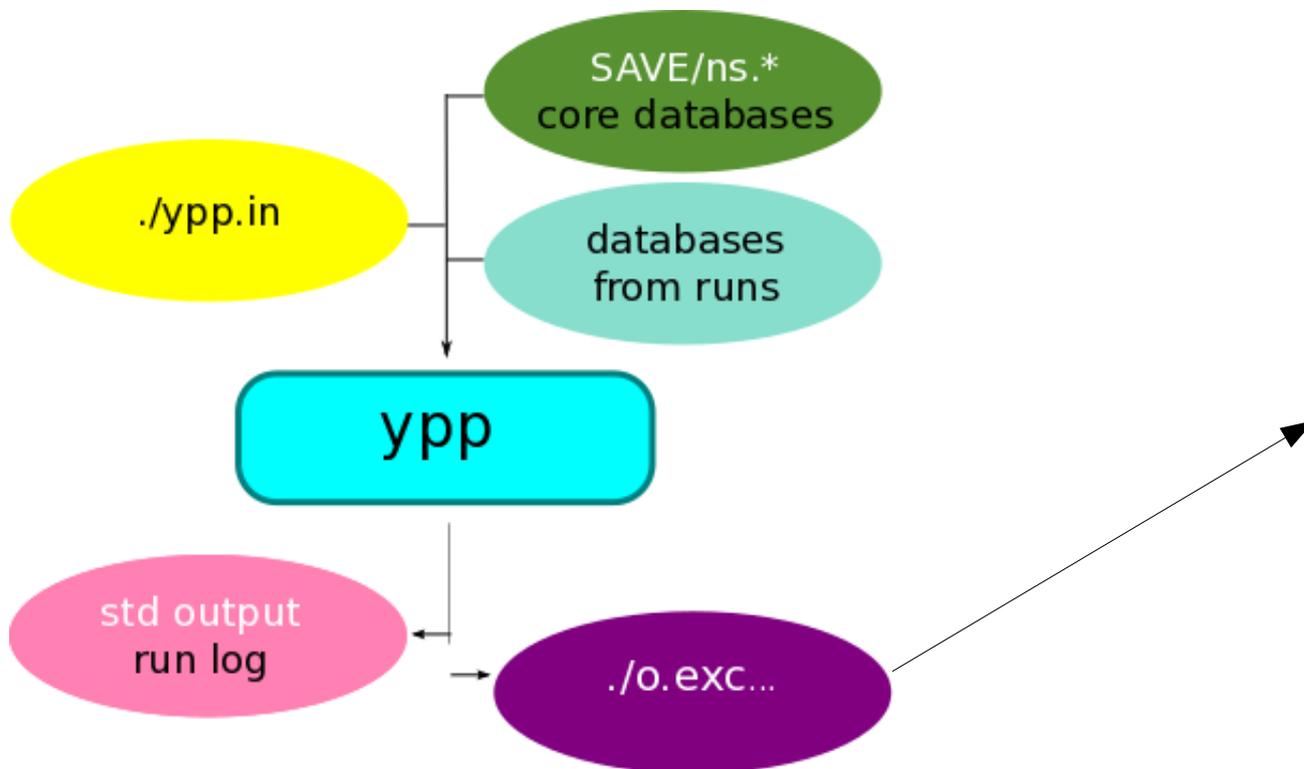
4. Run yambo



How it works



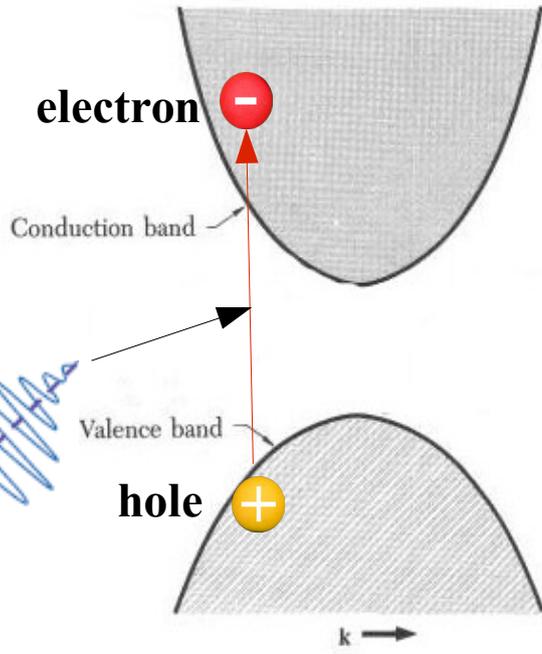
5. Post-processing with ypp



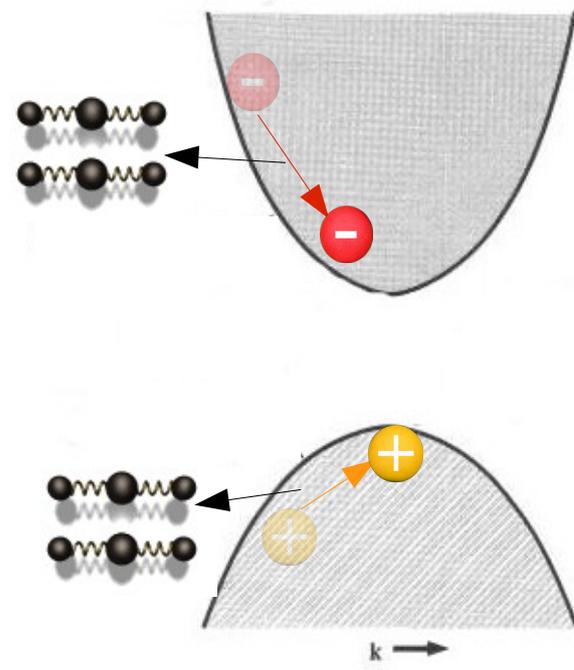
outline

- ❑ Motivations and history
- ❑ The code: **the structure of yambo**
- ❑ Applications: **few examples**
- ❑ Users view: **how it works in practice**
- ❑ Future: **new developments and more**

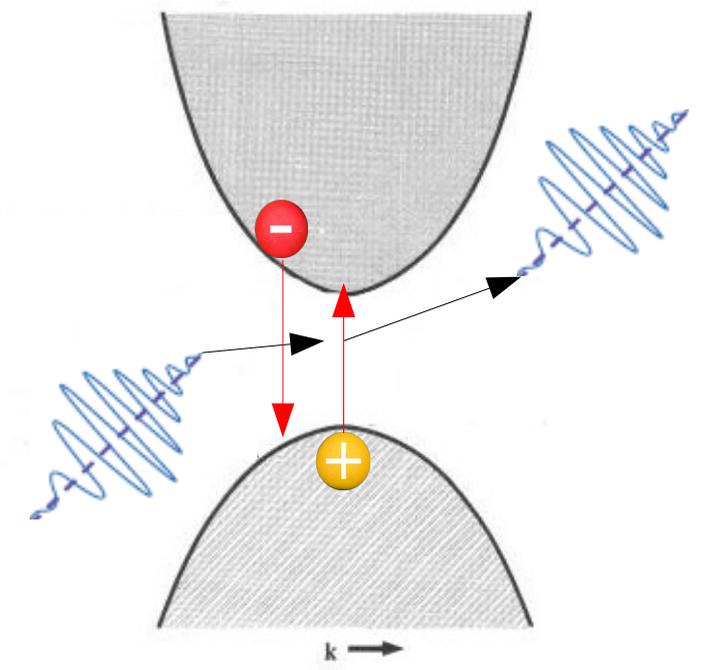
Non-equilibrium physics



1 – Photo-excitation process

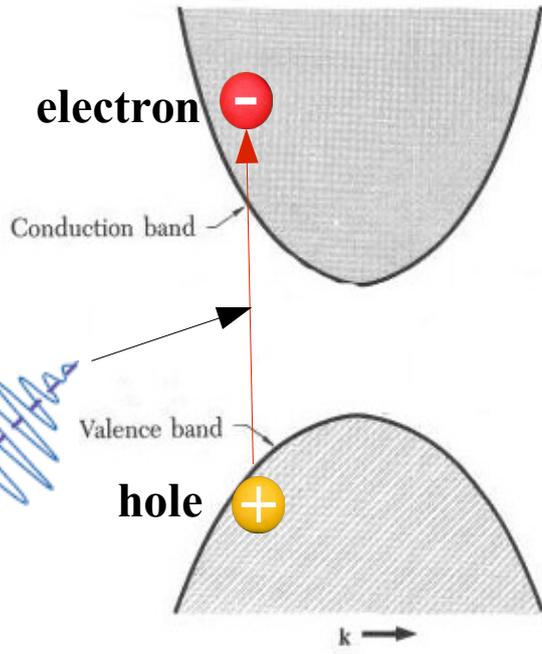


2 – relaxation

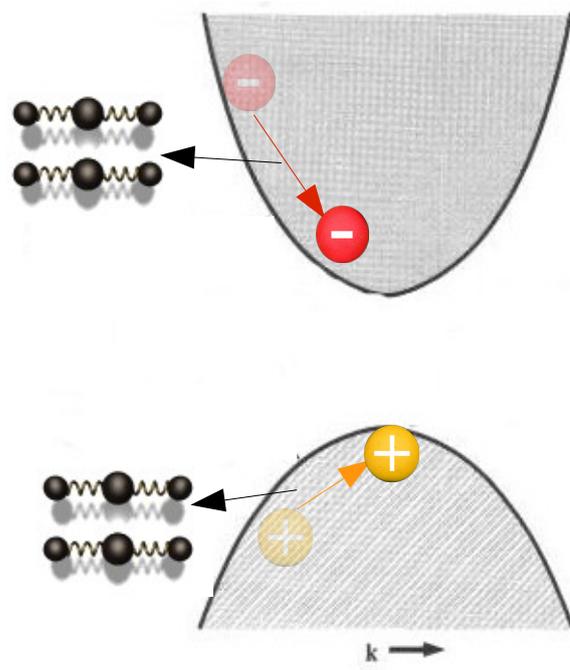


3 – Probing
(Transient absorption)

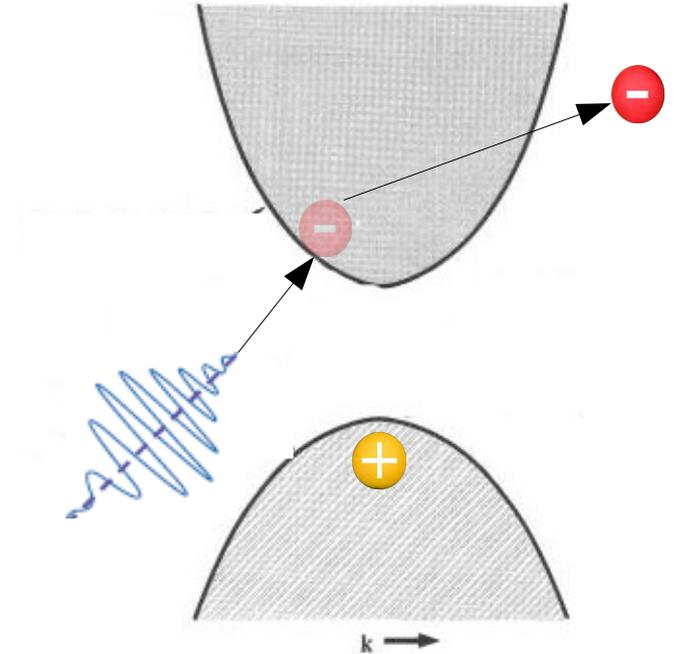
Non-equilibrium physics



1 – Photo-excitation process

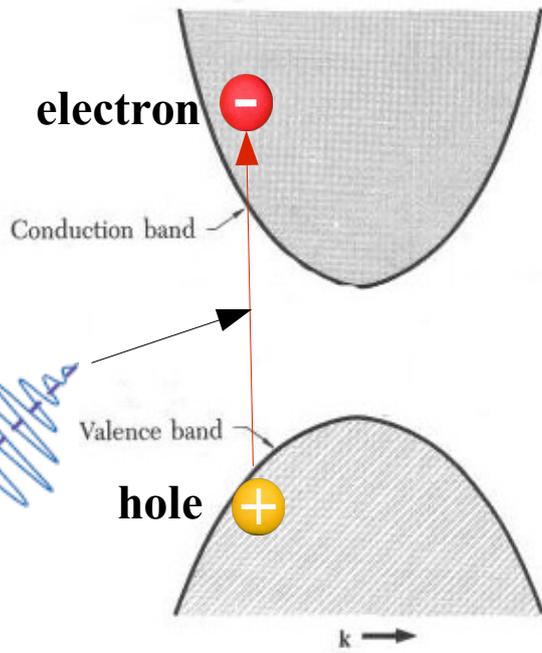


2 – relaxation

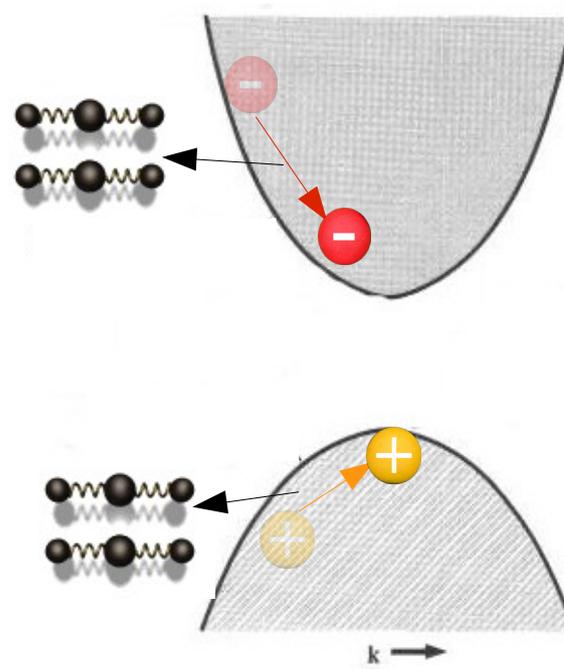


3 – Probing
(photo-emission)

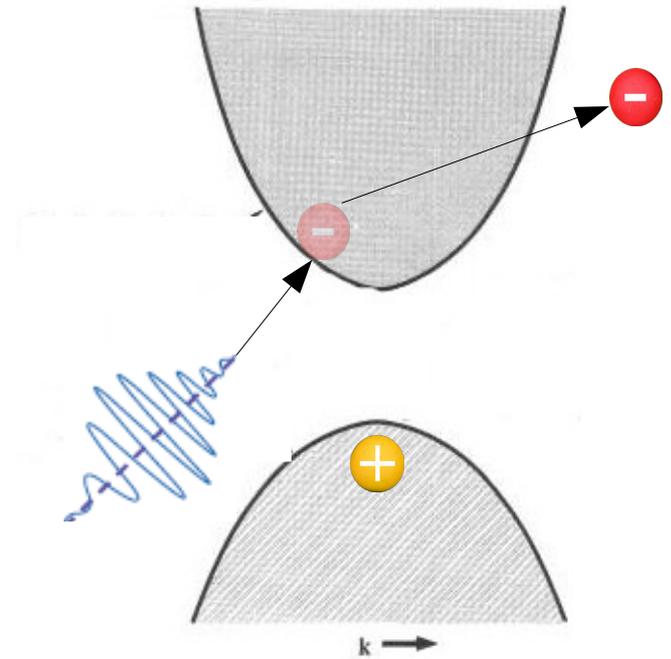
Non-equilibrium physics



1 – Photo-excitation process



2 – relaxation



3 – Probing
(photo-emission)

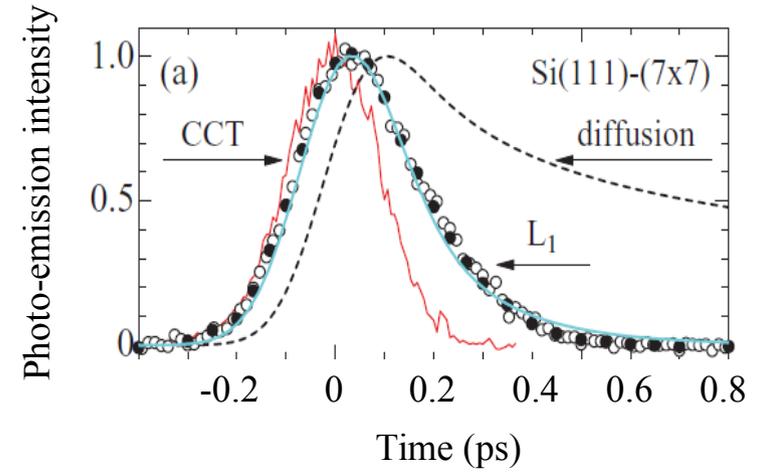
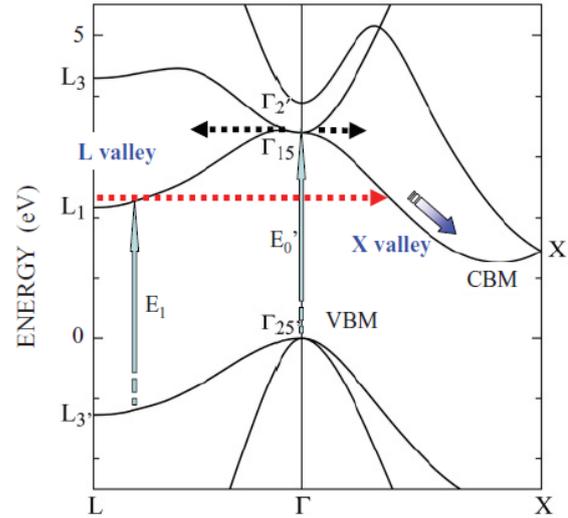
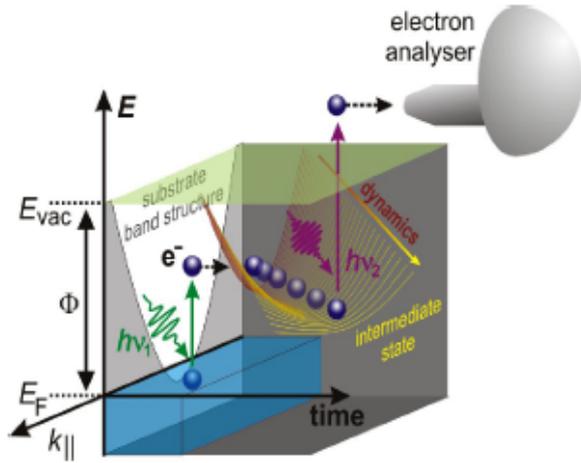
Propagate the equation of motion of the NEQ Green function

$$i \partial_t G_{nmk}^<(t, t) - [H_{QP}^{eq} + \Delta V^H + \Delta \Sigma_{xc}^{st} + U^{ext}(t), G^<(t, t)]_{nmk} = S_{nmk}(t)$$

GKBA approximation and more

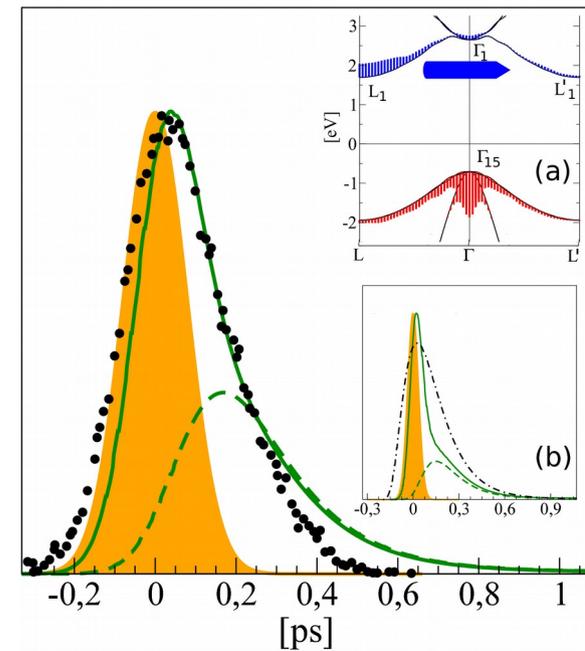
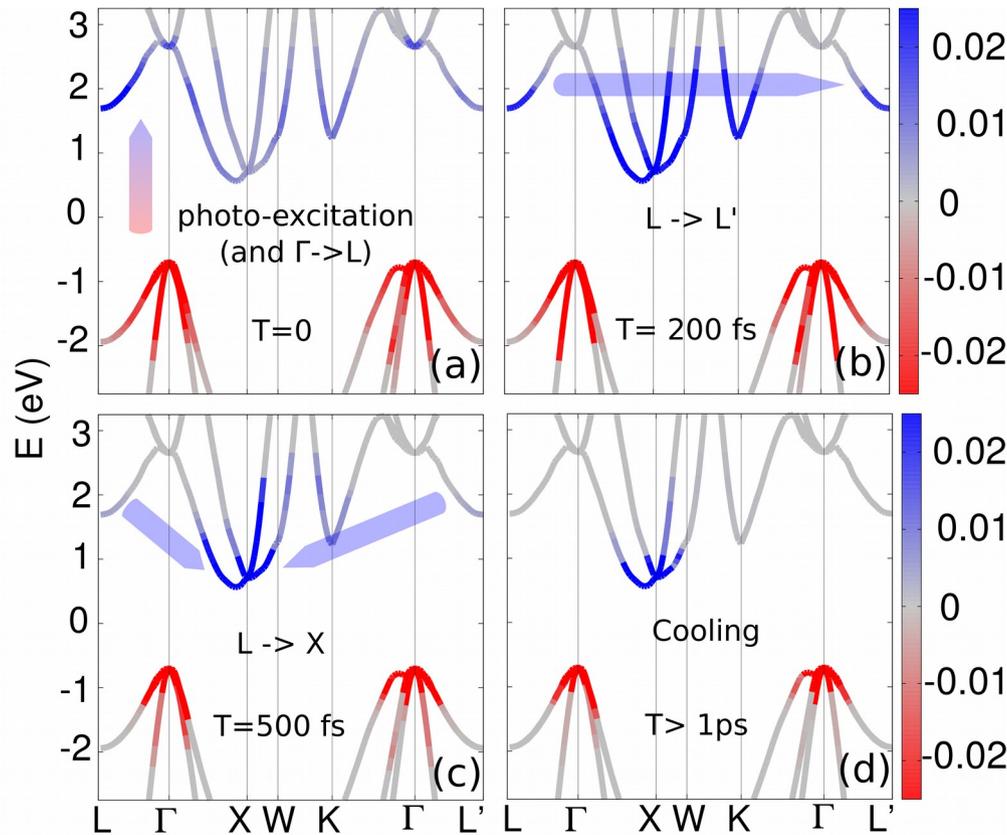
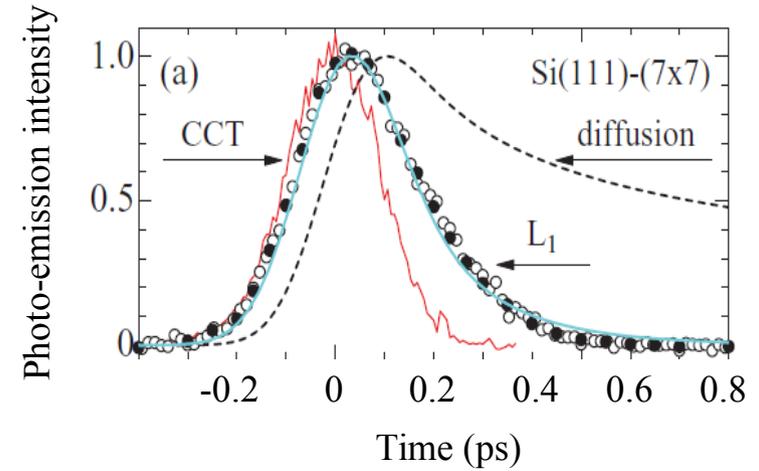
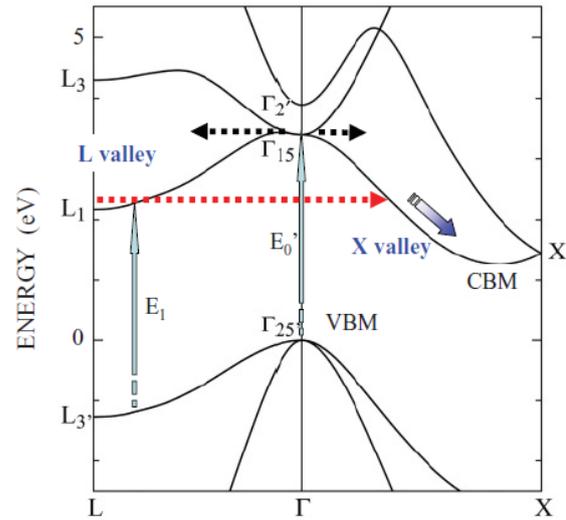
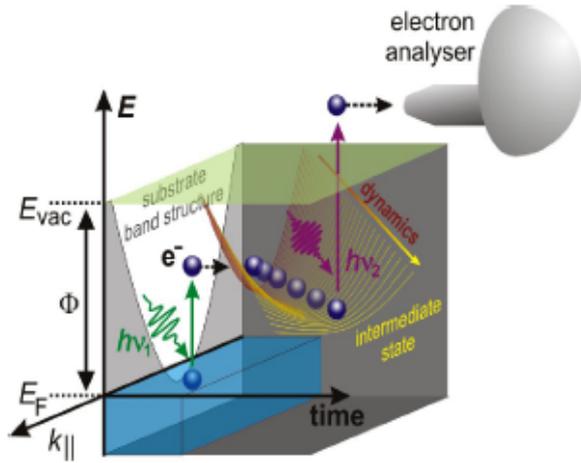
Not yet
GPL

Non-equilibrium physics



Not yet
GPL

Non-equilibrium physics



D. Sangalli, and A. Marini,
Europhysics Letters 110, 47004 (2015)



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Istituto di Struttura
della Materia



Thank you for your attention



[*www.yambo-code.org*](http://www.yambo-code.org)



[*en.wikipedia.org/wiki/YAMBO_code*](http://en.wikipedia.org/wiki/YAMBO_code)



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[*github.com/henriquemiranda/yambopy*](https://github.com/henriquemiranda/yambopy)

